

SOLUTIONS #2

QUESTION #1 PROBLEM 3.1

$\vec{B} = B_0 \hat{z}$, $\vec{U} = U_E \hat{x}$, POSITIVELY CHARGED PARTICLE

a) $\vec{U} = \vec{E} \times \vec{B} / q$, so $\vec{B} \times \vec{U} = \vec{E}$, AND $\vec{E} = (\hat{z} \times \hat{x}) B_0 U_E = \hat{y} B_0 U_E$

b) At $t=0$, $(x, y, z) = (0, 0, 0)$ AND $(v_x, v_y, v_z) = (0, 0, 0)$

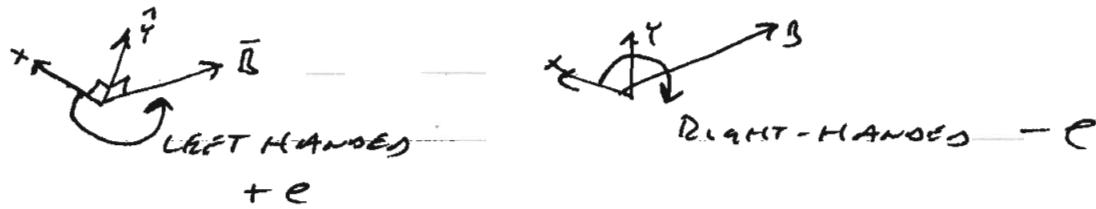
THEN, $v_z(t) = 0$ $z(t) = 0$

$$v_y(t) = -\rho \omega_c \cos(\omega_c t + \varphi) \quad y(t) = y_0 - \rho \sin(\omega_c t + \varphi)$$

$$v_x(t) = U_E - \rho \omega_c \sin(\omega_c t + \varphi) \quad x(t) = U_E t + \rho \cos(\omega_c t + \varphi)$$

WHERE $\omega_c = e B / m$

NOTE: FOR A POSITIVE CHARGE, CYCLOTHON MOTION IS "LEFT-HANDED" !!



QUESTION: WHAT ARE y_0 , ρ , AND φ ?

ANSWER: MATCH BOUNDARY/INITIAL VALUES...

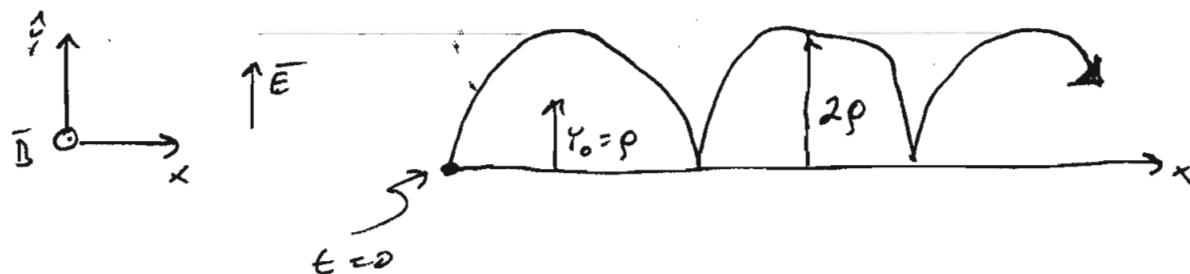
At $t=0$, $x=0 = \cos(\varphi)$

$$y=0 = y_0 - \rho \sin(\varphi)$$

$$v_x=0 = U_E - \rho \omega_c \sin(\varphi)$$

$$v_y=0 = \cos(\varphi)$$

Thus, $\varphi = \frac{\pi}{2}$, $y_0 = \rho$, AND $\varphi = U_E / \omega_c$



(2)

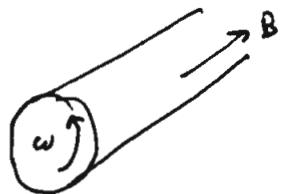
QUESTION #1 (CONTINUED)

c) Energy = K.E. + P.E.

$$\text{K.E.} = \frac{1}{2} m (v_x^2 + v_y^2) = \frac{1}{2} m (2v_E^2 - 2v_E^2 \sin(\omega_c t + \frac{\pi}{2})) \\ = m v_E^2 (1 - \sin(\omega_c t + \frac{\pi}{2}))$$

$$\text{P.E.} = -e v_E B_0 \gamma = -e v_E B_0 (r_0 - \rho \sin(\omega_c t + \frac{\pi}{2})) \\ = -\frac{e v_E^2 B_0}{\omega_c} (1 - \sin(\omega_c t + \frac{\pi}{2})) \\ = -m v_E^2 (1 - \sin(\omega_c t + \frac{\pi}{2}))$$

THUS, TOTAL ENERGY = 0

QUESTION #2 PROBLEM 3.2

a)

$$\bar{v}_E = \omega_0 r (\hat{z} \times \hat{r}) = \frac{\bar{E} \times \bar{B}}{B^2} = \frac{E_1}{B} (\hat{z} \times \hat{r})$$

$$\therefore E_R = -\omega_0 r B_0$$

b) $\rho = \epsilon_0 \nabla \cdot \bar{E} = -\epsilon_0 \frac{\omega_0 B_0}{r} \frac{2}{2r} (r^2) = -2\epsilon_0 \omega_0 B_0$

c) $\bar{E} = -\nabla \Phi \quad \therefore \frac{2\Phi}{2r} = \omega_0 B_0 r \Rightarrow \Phi(r) = \omega_0 B_0 \frac{r^2}{2}$

d) IF THERE WERE A UNIFORM CHARGE DENSITY, THEN THE COLUMN ROTATES AS A RIGID ROTOR. IMPOSING ROTATION, IN GENERAL, INVOLVES PLASMA TRANSPORT MODES AND BOUNDARY/EDGE EFFECTS.

(3)

Question #3 Problem 3.4

a) TWO METHODS TO DETERMINE A FIELD-LINE

METHOD #1

$$\frac{dn}{B_r} = \frac{n d\theta}{B_\theta} = \frac{ds}{|\theta|}$$

THUS

$$\int \frac{dn}{n} = 2 \int \frac{\cos \theta d\theta}{\sin \theta} = 2 \int \frac{d(\sin \theta)}{\sin \theta}$$

$$\ln n = 2 \ln(\sin \theta)$$

$$\text{ON } n \approx R \sin^2 \theta$$

METHOD #2

$$\bar{B} = \nabla \varphi \times \nabla \psi \quad \text{so } (\psi, \varphi) \text{ LABEL A FIELD-LINE}$$

since $\nabla \varphi = \frac{\hat{\phi}}{r \sin \theta}$, we write

$$\begin{aligned} \hat{\theta} \cdot \bar{B} &= \hat{\theta} \cdot (\nabla \varphi \times \nabla \psi) = (\hat{\theta} \times \hat{\varphi}) \cdot \frac{1}{r \sin \theta} \left(\frac{2\psi}{2\theta} \hat{r} + \frac{1}{r \sin \theta} \frac{2\psi}{2\theta} \hat{\theta} \right) \\ &= \frac{1}{r \sin \theta} \frac{2\psi}{2\theta} \\ &= \frac{\mu_0 M}{4\pi r^2} \sin \theta \end{aligned}$$

SOLVING FOR $\psi(r)$...

$$\psi(r, \theta) = -\frac{\mu_0 M}{4\pi r} \sin^2 \theta$$

so $\frac{\sin^2 \theta}{r} = \text{constant along a field line}$

$$b) \bar{F} = (\bar{B} \cdot \bar{v}) \hat{b} \propto \frac{1}{R_c} \quad \hat{b} = \frac{2 \cos \theta \hat{r} + \sin \theta \hat{\theta}}{3 \cos^2 \theta + 1}$$

in spherical coordinates

$$\begin{aligned} \hat{r} \cdot \bar{F} \Big|_{\theta \rightarrow \frac{\pi}{2}} &= \frac{\hat{b}_\theta}{r} \frac{2b_r}{2\theta} - \frac{\hat{b}_\theta}{r} = \frac{1}{r} \frac{2}{2\theta} \left(\frac{2 \cos \theta}{3 \cos^2 \theta + 1} \right) \Big|_{\theta=\frac{\pi}{2}} - \frac{1}{r} \\ &= \frac{1}{r} [-2 - 1] = -\frac{3}{r} \quad \text{QED} \end{aligned}$$

PROBLEM #3.4 (CONTINUED)

NOTE ALSO $\nabla B = \hat{z} \frac{\partial B}{\partial r} + \hat{\theta} \frac{\partial B}{\partial \theta}$

$$\frac{\partial B}{\partial r} = -\frac{3B}{r}. \text{ THEREFORE } \hat{z} \cdot \vec{F} = \hat{z} \cdot \nabla B / B$$

c) $\vec{V}_x = \frac{2w_{||}}{8B} \left[\frac{\hat{b} \times (-\hat{z} R_c)}{|R_c|^2} \right] = \frac{2w_{||}}{8B} \left[\hat{b} \times \vec{F} \right]$

$$= \frac{2w_{||}}{8B} \left(\hat{b} \times \left(-\hat{z} \frac{3}{r} \right) \right) \text{ At } \theta = \frac{\pi}{2}$$

$$= \hat{\phi} \frac{6w_{||}}{8B_1}$$

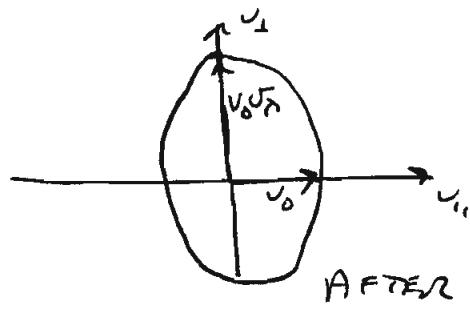
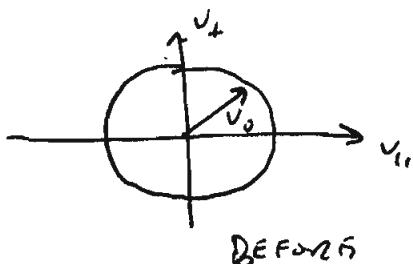
d) $\vec{V}_{\nabla B} = \frac{w_{\perp}}{8B} \left[\hat{b} \times \frac{\nabla B}{B} \right] = \hat{\phi} \frac{3w_{\perp}}{8B_1} \text{ @ } \theta = \frac{\pi}{2}$

e) $\frac{V_x}{V_{\nabla B}} = \frac{2w_{||}}{w_{\perp}}$

QUESTION #4 THIS IS A VELOCITY-SPACE PROBLEM
 RECOGNIZING THAT $\mu = \frac{1}{2}v^2/B = \text{constant}$ DURING
 ADIABATIC CHANGES IN THE MAGNETIC FIELD.

FOR AN ADIABATIC INCREASE IN B , THEN

$$v_{\perp}' \rightarrow \sqrt{\lambda} v_{\perp}.$$



QUESTION H4 (CONT.)

In order to solve this problem, it is useful to practice integrating velocity-space distribution functions. For simplicity, we'll normalize the density and particle mass to 1: $m = m = 1$. Then, the distribution function has the following properties

$$\iiint d^3v f = 1$$

$$\iiint d^3v \frac{1}{2} v^2 f = \langle \frac{1}{2} v^2 \rangle = \text{K.E.}$$

$$\iiint d^3v \frac{1}{2} v_{\perp}^2 f = \langle \frac{1}{2} v_{\perp}^2 \rangle = \text{PERPENDICULAR ENERGY}$$

$$\iiint d^3v \frac{1}{2} v_{\parallel}^2 f = \langle \frac{1}{2} v_{\parallel}^2 \rangle = \text{PARALLEL ENERGY.}$$

LET'S LOOK AT THE INITIAL DISTRIBUTION FUNCTION

$$f(v) = N \delta(v - v_0)$$

WHERE N = NORMALIZATION

and $\delta(\dots)$ is the famous "DIRAC" DELTA FUNCTION

THIS DISTRIBUTION IS A SPHERE OF RADIUS v_0 IN VELOCITY SPACE. THE NORMALIZATION IS

$$\begin{aligned} \iiint d^3v f = 1 &= \int_0^\infty 4\pi v^2 dv N \delta(v - v_0) \\ &= 4\pi v_0^2 N \quad \therefore \quad N = \frac{1}{4\pi v_0^2} \end{aligned}$$

WE ALSO HAVE $\langle \frac{1}{2} v^2 \rangle = \frac{1}{2} v_0^2$ AS EXPECTED.

TO FIND THE PERPENDICULAR KINETIC ENERGY, WE WRITE IN CYLINDRICAL VELOCITY-SPACE

$$\iiint d^3v \frac{1}{2} v_{\perp}^2 f = \langle \frac{1}{2} v_{\perp}^2 \rangle = \int_0^\infty 2\pi v_{\perp} dv_{\perp} \int_{-\infty}^\infty dv_{\parallel} \frac{1}{2} v_{\perp}^2 \frac{\delta(\sqrt{v_{\perp}^2 + v_{\parallel}^2} - v_0)}{4\pi v_0^2}$$

QUESTION #4 (cont.)

REMEMBER THE PROPERTY OF DIRAC DELTA FUNCTIONS!

$$\int_{-\infty}^{\infty} dx h(x) \delta(g(x)) = \int_{-\infty}^{\infty} dx h(x) \delta(g'(x-x_0)) \text{ where } g(x_0)=0$$

$$= \frac{h(x_0)}{|g'|} \quad g' = \left. \frac{dg}{dx} \right|_{x=x_0}$$

THEREFORE,

$$\langle \frac{1}{2} v_{\perp}^2 \rangle = \int_{-v_0}^{+v_0} dv_{\parallel} \frac{1}{2v_0} \frac{1}{2} (v_0^2 - v_{\parallel}^2) = \frac{2}{3} \left(\frac{1}{2} v_0^2 \right)$$

$$\text{WE ALSO HAVE } \langle \frac{1}{2} v_{\parallel}^2 \rangle = \int_{-v_0}^{v_0} dv_{\parallel} \frac{1}{2v_0} \frac{1}{2} v_{\parallel}^2 = \frac{1}{3} \left(\frac{1}{2} v_0^2 \right)$$

$$\text{NOTICE } \langle \frac{1}{2} v^2 \rangle = \langle \frac{1}{2} v_{\perp}^2 \rangle + \langle \frac{1}{2} v_{\parallel}^2 \rangle \text{ AS EXPECTED.}$$

Now, we can compute what happens after an adiabatic compression where $v_{\perp} \rightarrow \sqrt{\lambda} v_{\perp}$.

THE DISTRIBUTION HAS THE FORM

$$f(v) = n \delta\left(\sqrt{\frac{v_{\perp}^2}{\lambda} + v_{\parallel}^2} - v_0\right)$$

FIRST, WE NORMALIZE:

$$1 = \int_0^{\infty} 2\pi v_{\perp} dv_{\perp} \int_{-\infty}^{\infty} dv_{\parallel} n \delta\left(\sqrt{\frac{v_{\perp}^2}{\lambda} + v_{\parallel}^2} - v_0\right)$$

$$= 4\pi v_0^2 \lambda n \therefore n = \frac{1}{4\pi v_0^2 \lambda}$$

ON THE NEXT PAGE, WE COMPUTE $\langle \frac{1}{2} v_{\perp}^2 \rangle$

QUESTION #4 (CONT.)

$$\text{K.E. (AFTER COMPRESSION)} = \int_0^\infty 2\pi v dy_1 \int_{-\infty}^\infty dv_1 \frac{1}{2} (v_1^2 + v_2^2) f$$

$$= \frac{1}{2v_0^2 \lambda} \int_0^\infty v_1 dy_1 \int_{-\infty}^\infty dv_1 \frac{1}{2} (v_1^2 + v_2^2) \delta(\sqrt{\frac{y_1^2}{\lambda} + v_1^2} - v_0)$$

LET'S CHANGE VARIABLES: $\tilde{v}_1^2 = \frac{y_1^2}{\lambda}$ so $v_1 dy_1 = \lambda d\tilde{v}_1 \tilde{v}_1$

$$\text{K.E.} = \frac{1}{2v_0^2} \int_0^\infty \tilde{v}_1 d\tilde{v}_1 \int_{-\infty}^\infty dv_1 \frac{1}{2} (\lambda \tilde{v}_1^2 + v_2^2) \delta(\sqrt{\tilde{v}_1^2 + v_2^2} - v_0)$$

$$= \frac{1}{2v_0^2} \int_0^\infty \tilde{v}_1 d\tilde{v}_1 \int_{-\infty}^\infty dv_1 \left(\frac{1}{2} v^2 + (\lambda - 1) \frac{1}{2} \tilde{v}_1^2 \right) \delta(\sqrt{\tilde{v}_1^2 + v_2^2} - v_0)$$

$$= \left\langle \frac{1}{2} v^2 \right\rangle + (\lambda - 1) \left\langle \frac{1}{2} \tilde{v}_1^2 \right\rangle = \frac{1}{2} v_0^2 + (\lambda - 1) \frac{2}{3} \left(\frac{1}{2} v_0^2 \right)$$

THE REST IS EASY!

!!!

0. INITIAL K.E. = $\frac{1}{2} v_0^2$

1. AFTER ADIABATIC COMPRESSION, NEW K.E. IS
 $K.E. \rightarrow \frac{1}{2} v_0^2 \left[1 + \frac{2}{3} (\lambda - 1) \right]$

2. AFTER VELOCITY DISTRIBUTION BECOMES ISOTROPIC,
 THE K.E. DOES NOT CHANGE

3. AFTER ADIABATIC DE-COMPRESSION NEW K.E. IS
 $K.E. \rightarrow \frac{1}{2} v_0^2 \left[1 + \frac{2}{3} (\lambda - 1) \right] \left[1 + \frac{2}{3} \left(\frac{1}{\lambda} - 1 \right) \right]$

4. AFTER ISOTROPIC, K.E. REMAINS UNCHANGED
 THE FINAL K.E. IS

$$K.E. (\text{FINAL}) = \frac{1}{2} v_0^2 \left[1 + \frac{2}{3} (\lambda - 1) \right] \left[1 - \frac{2}{3} \frac{(\lambda - 1)}{\lambda} \right]$$

$$= \frac{1}{2} v_0^2 \left[1 + \frac{2}{9} \frac{(\lambda - 1)^2}{\lambda} \right]$$

FINALLY, THE MAXIMUM ENERGY

IS $E_{\max} = \frac{1}{2} v_0^2 \lambda$ THE MINIMUM ENERGY IS $\frac{1}{2} v_0^2 \frac{1}{\lambda}$.

NOTICE THAT K.E. (FINAL) $\propto \lambda$ AS $\lambda \rightarrow \infty$.