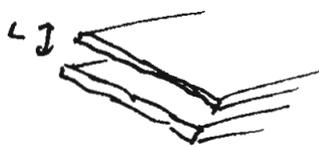


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SOLUTIONS: HW 4

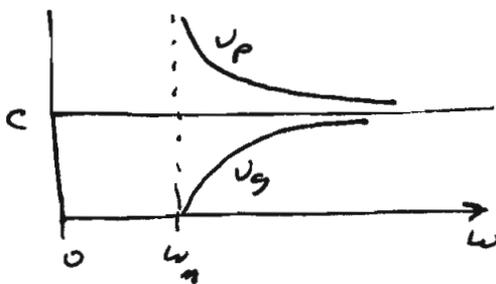
#1] p. 4.3 

$$\omega^2 = \omega_m^2 + c^2 k^2$$

$$v_p = \frac{\omega}{k} = c \frac{\omega}{c k} = \frac{c}{\sqrt{1 - \omega_m^2/\omega^2}}$$

$$v_g = \frac{2\omega}{2k}, \text{ BUT } 2\omega \frac{2\omega}{2k} = 2c^2 k$$

$$\text{SO } v_g = c^2 k / \omega = c^2 / v_p = c \sqrt{1 - \omega_m^2/\omega^2}$$



THIS IS JUST LIKE E.M.

PLASMA WAVES, WHEN

$$\omega_m^2 = \omega_p^2$$

#2] p. 4.4

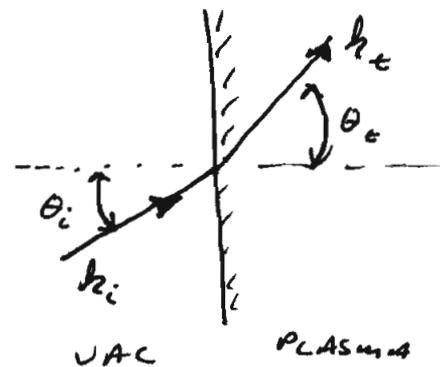
SNELL'S LAW:

$$n_i \sin \theta_i = n_e \sin \theta_e$$

WHEN $\theta_e = \pi/2$, WE HAVE

"TOTAL REFLECTION". THUS,

THE CRITICAL ANGLE IS



$$\sin \theta_c = n_e / n_i . \quad n_i = 1 . \quad n_e = \frac{c k}{\omega} = \sqrt{1 - \omega_p^2 / \omega^2}$$

$$\text{SO } \sin \theta_c = \sqrt{1 - \omega_p^2 / \omega^2}$$

#3] p. 4.5 FOR HARMONIC PERTURBATIONS, $E(t) = \text{Re} \{ E_0 e^{-i\omega t} \}$

$$\text{SO... } -i m_s \omega v_s + v_s m_s \omega v_s = e_s (E + v_s \times B_0)$$

$$-i \omega m_s \left(1 + \frac{v_s}{\omega} \right) v_s = e_s (E + v_s \times B_0)$$

$$m_s^* = m_s \left(1 + \frac{v_s}{\omega} \right)$$

#3) CONTINUED

b) FOR FIELD-FREE E, M WAVES

$$k^2 c^2 = \omega^2 - \omega_p^2$$

$$\text{So } \frac{kc}{\omega} = \sqrt{1 - \frac{\omega_p^2}{\omega^2} \left(1 + i \frac{v_s}{\omega}\right)^{-1}}$$

IF $\omega_p^2 \ll \omega^2$ AND $v_s \ll \omega$, THEN

$$\begin{aligned} \frac{kc}{\omega} &\approx 1 - \frac{1}{2} \left(\frac{\omega_p^2}{\omega^2}\right) \left(1 - i \frac{v_s}{\omega} + \dots\right) + \dots \\ &\approx 1 - \frac{1}{2} \left(\frac{\omega_p^2}{\omega^2}\right) + i \frac{1}{2} \frac{\omega_p^2 v_s}{\omega^3} + \dots \end{aligned}$$

c) FOR LONGITUDINAL WAVES ...

$$-i \omega m_s \left(1 + i \frac{v_s}{\omega}\right) v_s = e_s E$$

$$\text{So } \omega^2 = \omega_p^2 \left(1 + i \frac{v_s}{\omega}\right)^{-1}$$

$$\begin{aligned} \text{AND } \omega &\approx \omega_p \left(1 - \frac{1}{2} i \frac{v_s}{\omega} + \dots\right) \\ &\approx \omega_p - \frac{i}{2} v_s \end{aligned}$$

SO ...

$$\begin{aligned} E(t) &= E_0 \left\{ e^{-i\omega_p t} e^{-v_s t/2} \right\} \\ &= E_0 \cos(\omega_p t) e^{-v_s t/2} \end{aligned}$$

#4) P. 4.7

$$\text{Det} \begin{bmatrix} S - m^2 \cos^2 \theta & -iD & m^2 \sin \theta \cos \theta \\ iD & S - m^2 & 0 \\ m^2 \sin \theta \cos \theta & 0 & P - m^2 \sin^2 \theta \end{bmatrix}$$

$$\begin{aligned} \text{Det} &= (S - m^2 \cos^2 \theta)(S - m^2)(P - m^2 \sin^2 \theta) \\ &\quad - m^4 \sin^2 \theta \cos^2 \theta (S - m^2) - D^2 (P - m^2 \sin^2 \theta) \\ &= (S - m^2) \left[SP - m^2 \{ P \cos^2 \theta + S \sin^2 \theta \} + m^4 \sin^2 \theta \cos^2 \theta \right. \\ &\quad \left. - m^4 \sin^2 \theta \cos^2 \theta \right] - D^2 (P - m^2 \sin^2 \theta) \\ &= m^4 (P \cos^2 \theta + S \sin^2 \theta) - m^2 [SP + S(P \cos^2 \theta + S \sin^2 \theta) - D^2 \sin^2 \theta] \\ &\quad + S^2 P - D^2 P \end{aligned}$$

#4) (CONT.)

$\tau_0 \text{ DRT} = A m^4 - B m^2 + C$

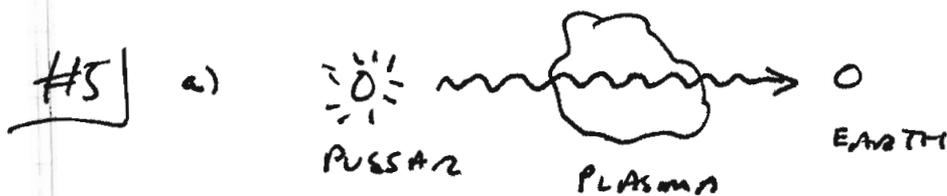
$A = S \sin^2 \theta + P \cos^2 \theta$

$B = SP(1 + \cos^2 \theta) + S^2 \sin^2 \theta - D^2 \sin^2 \theta$

$C = P(S^2 - D^2) = PRL$

BUT $S^2 - D^2 = RL$ (see P. 4.6)

AND $B = SP(1 + \cos^2 \theta) + RL \sin^2 \theta$ $\sigma \in \sigma$



$\tau = \int \frac{ds}{v_g}$

$v_g = \frac{2\omega}{2\hbar}$

$\omega^2 = k^2 c^2 + \omega_p^2$

$\frac{2\omega}{2\hbar} = \frac{c^2 \hbar}{\omega} = \frac{c}{v} \sqrt{\omega^2 - \omega_p^2}$

\therefore IF $\frac{\omega_p^2}{\omega^2} \ll 1$, THEN

$\tau = \int \frac{ds}{c} \left(1 - \frac{\omega_p^2}{\omega^2}\right)^{-\frac{1}{2}} \approx \int \frac{ds}{c} \left(1 + \frac{1}{2} \frac{\omega_p^2}{\omega^2} + \dots\right)$

$\approx \int \frac{ds}{c} + \frac{1}{2c\omega^2} \int ds \omega_p^2$

$\omega_p^2 = \frac{e^2 n}{m_0 \epsilon_0}$

$\approx \int \frac{ds}{c} + \frac{e^2}{2c m_0 \epsilon_0 \omega^2} \int ds n$

b) COEFFICIENT, $C = \frac{e^2}{2c m_0 \epsilon_0 (2\pi)^2} = \frac{c}{2\pi} \left(\frac{e^2}{4\pi \epsilon_0 m_0 c^2} \right)$

CLASSICAL RADIUS OF ELECTRON
 $2.8 \times 10^{-15} \text{ m}$

c) $\tau_{100m} - \tau_{200m} = 2 = D \left[\frac{1}{(10^8)^2} - \frac{1}{4(10^8)^2} \right] = \frac{3}{4} \frac{D}{10^{16}}$

$\therefore D = \frac{8}{3} \times 10^{16}$

(4)

#5 | CONT.

D) IF $n = 0.03 \text{ cm}^{-3} = 3 \times 10^4 \text{ pm}^{-3}$, THEN

$$D = \frac{2}{24} 2.8 \times 10^{-15} \text{ L } 3 \times 10^4 = \frac{8}{5} \times 10^{16}$$

$$\therefore L = \frac{8}{5} \times 10^{16} \frac{2\pi}{c} \frac{1}{3 \times 10^4} \frac{1}{2.8 \times 10^{-15}}$$

$$= 6.6 \times 10^{18} \text{ m}$$

BUT $1 \text{ PARSEC} = 3 \times 10^{16} \text{ m}$ SO $L = 222 \text{ PARSEC}$ E) MAGNETIC FIELDS, REFRACTION, AND SCATTERING
MAKE THIS MEASUREMENT DIFFICULT.